

HUVINETT


# Radioactivity

Course on Nuclear Fundamentals  
2<sup>nd</sup> lecture

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- Radioactive decay
- Alpha, beta, gamma-decay,
- Exponential decay law, half-life, activity
- Poisson distribution
- Radioactive decay chains
- Radioactive dating
- Self-test questions

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## Alpha, beta, gamma-decay

Reminder: the average energy of a nucleon:  
Radioactive decay: **spontaneous** process

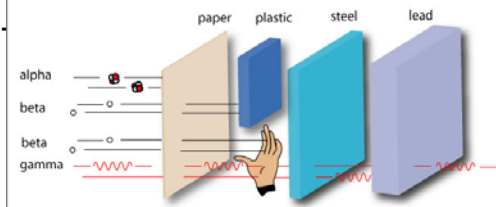
It can only occur if  $\varepsilon_{\text{final}} < \varepsilon_{\text{initial}}$

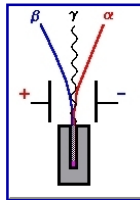
The energy released = (mass difference)· $c^2$

Three types of particles are emitted:

- $\alpha$  – particles:  ${}^4_2\text{He}$  nuclei
- $\beta$  – particles: high energy **electrons/positrons**
- $\gamma$  - radiation: **electromagnetic** (photons)

The penetration distance (range) are different:





$$\varepsilon = \frac{E}{A} = -\frac{B}{A}$$

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During radioactive decays:  
nuclei **transform** into each-other

**Parent** nucleus decays into **daughter** nucleus

### Conserved physical quantities

- **energy** (taking  $E = mc^2$  into account)
- **mass number** /number of nucleons/ ( $A$ )
- **electric charge** (nucleus:  $+Ze$ , electron:  $-e$ )
- **several others**  
(momentum, angular momentum, parity, etc.)

**We will consider only the three most important**

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### Gamma-decay

The **gamma-decay**: between energy levels of the nucleus  
 The **composition** does not change  
 (mass-number and electric charge conservation  
 is automatically fulfilled):

$$\boxed{{}_Z^A X^* \rightarrow {}_Z^A X + \gamma}$$

**Energy conservation:**  
 Gamma-transitions occur between the energy levels  
 Photon energy:  $h\nu = E_\gamma = E_1 - E_2$

Ground state  $0^+$   ${}_{28}^{60}\text{Ni}$  0,0 MeV

### Alpha-decay

The **alpha-decay**: an alpha-particle (He nuclei) is emitted:

Mass-number conservation

$$\boxed{{}_Z^A X \rightarrow {}_{Z-2}^{A-4} Y + {}_2^4 \text{He}}$$

Electric charge conservation

Large, unstable nucleus → Smaller, more stable nucleus + Alpha particle

Occurs only for heavy nuclei, because only for them will be energy released, if the nucleus gets smaller!

Average energy of a nucleon in function of mass number

### Beta-decays

Light particles (leptons) are emitted.

**Mass-number (A)** does not change (isobar transitions)  
 (Mass-number is conserved automatically)

From the Weizsäcker-formula we get ( $A = \text{const. case}$ )

During beta-decays the **atomic number (Z)** changes by  $\pm 1$ .

$\Delta Z = +1$   
 $\Delta Z = -1$

Electric charge conservation: since  $Z$  changes, the emitted (light) particles should carry electric charge

**Three types:**

Negative beta-decay ( $Z$  increasing)  ${}_Z^A X \rightarrow {}_{Z+1}^A Y + e^- + \bar{\nu}$

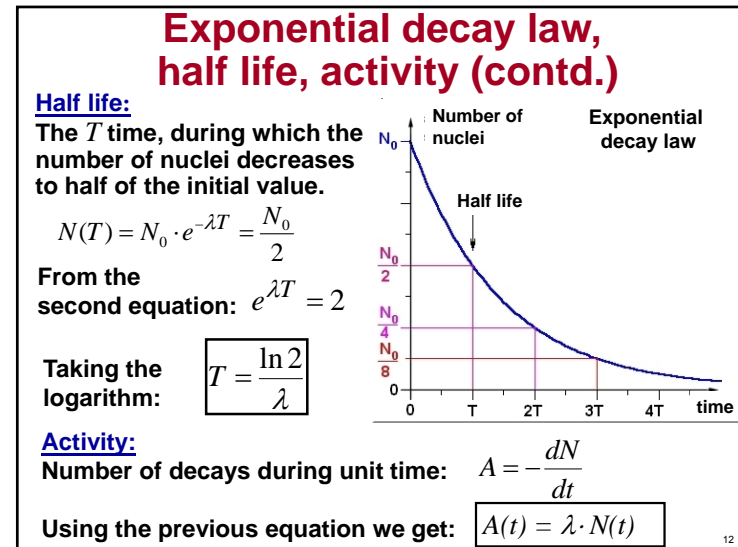
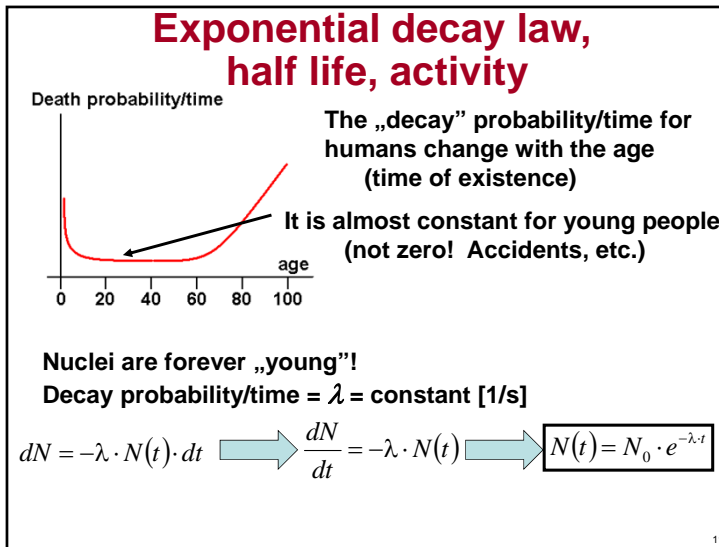
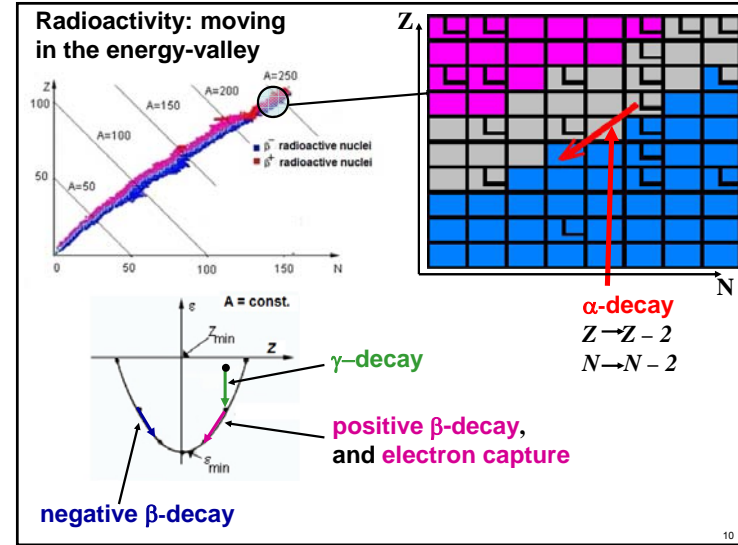
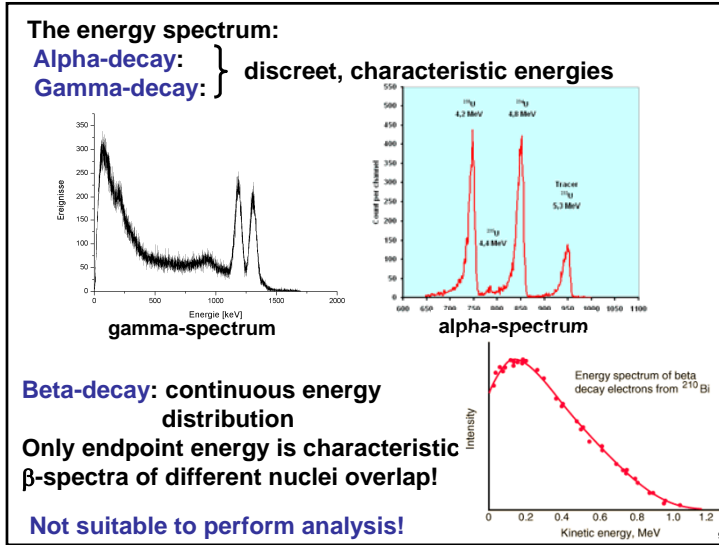
Positive beta-decay ( $Z$  decreasing)  ${}_Z^A X \rightarrow {}_{Z-1}^A Y + e^+ + \nu$

Electron capture ( $Z$  decreasing)  ${}_Z^A X + e^- \rightarrow {}_{Z-1}^A Y + \nu$

Here  $e^-$  : electron,  $\nu$  : neutrino,  $\bar{\nu}$  : antineutrino

Released energy is shared randomly } Energy spectrum of the electron (positron) and the antineutrino (neutrino) is continuous!

Negative beta-decay: when too much **neutrons**  
 Positive beta-decay: } when too much **protons** in the nucleus  
 Electron capture: }



## Poisson-distribution

The radioactive decay is a **stochastic (statistical)** process! (described by the decay **probability/time**:  $\lambda$ )

- The exact decay time of a **single** atom cannot be predicted
- The exponential decay law can be used only for **large number** of particles.

The **Poisson-distribution** gives the probability that

- exactly  $k$  decays occur
- in a source of  $a$  activity
- during  $t$  time.

$$P(k, at) = \frac{(at)^k}{k!} e^{-at}$$

Can be used only if the  $a$  activity can be considered constant ( $t \ll T_{half}$ ).

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## Poisson-distribution (contd.)

$$P(k, at) = \frac{(at)^k}{k!} e^{-at}$$

expectation value of  $k$ :  $\langle k \rangle = a \cdot t$

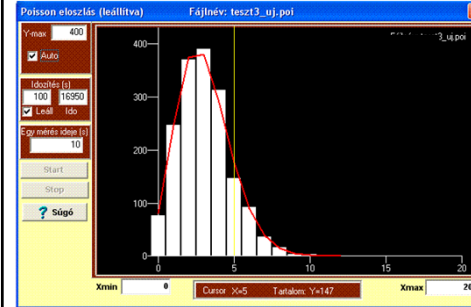
scatter of  $k$ :  $\sigma_k = \sqrt{a \cdot t}$

Please note:

$$\sigma_k = \sqrt{\langle k \rangle}$$

**Example:**

If  $N=100$  decays are expected, then the number of decays will be between 90 and 110 in most cases.



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## Radioactive decay chains

If heavy nuclei (with large  $A$ ) are decaying, in many cases also the daughter nuclei are radioactive. They decay further, until finally the chain terminates by reaching a stable nucleus.

Only alpha-decay changes the mass-number, and always decreases it by four.

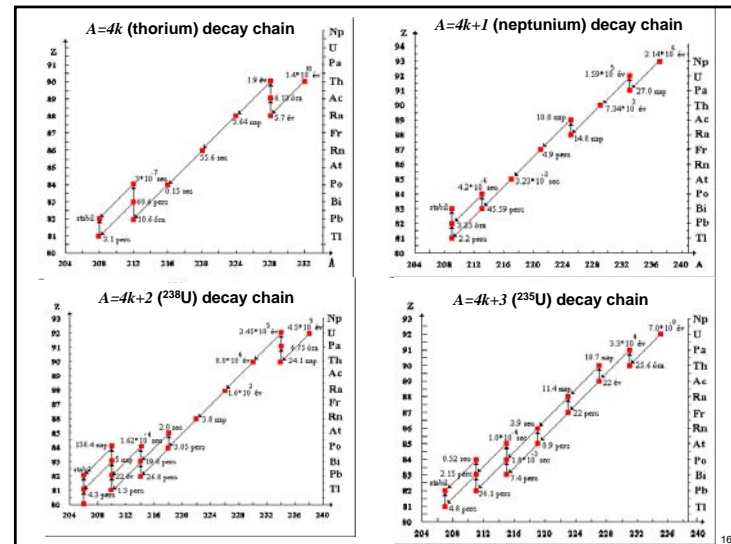
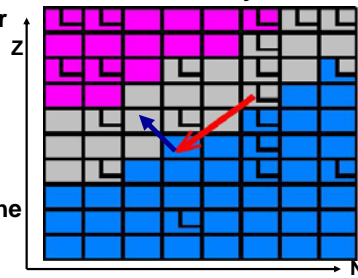
Consequence: the mass-number of every member of the same chain gives the same remainder if divided by 4.

Therefore there are only four decay chains:

$$A = 4k, A = 4k+1,$$

$$A = 4k+2, A = 4k+3$$

The  $\alpha$ -decays are followed by  $\beta$ -decays (and  $\gamma$ -decays), so that the chain can follow the incline of the energy valley.



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**Radioactive equilibrium**

Consider a radioactive „family” consisting only from 3 members:  $1 \rightarrow 2 \rightarrow 3$ , and denote the decay constants by  $\lambda_1$  and  $\lambda_2$ . The number of the nuclei:  $N_1(t)$ ,  $N_2(t)$ ,  $N_3(t)$ .

The equations describing the change in the number of nuclei:

$$\frac{dN_1}{dt} = -\lambda_1 \cdot N_1(t) \quad (\text{only decays})$$

$$\frac{dN_2}{dt} = -\lambda_2 \cdot N_2(t) + \lambda_1 \cdot N_1(t) \quad (\text{decays and is generated from the previous})$$

$$\frac{dN_3}{dt} = +\lambda_2 \cdot N_2 \quad (\text{only generated from the previous})$$

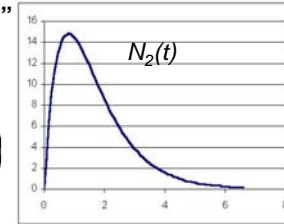
The solution of the first equation is already familiar:  $N_1(t) = N_1(0) \cdot e^{-\lambda_1 t}$

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Initial conditions:  $N_2(0) = 0$  and  $N_3(0) = 0$ , i.e. initially we have no „2” and „3” material, only material „1”.

The solution for  $N_2(t)$ :

$$N_2(t) = N_1(0) \frac{\lambda_1}{\lambda_2 - \lambda_1} \left( e^{-\lambda_1 t} - e^{-\lambda_2 t} \right) \quad (\text{if } \lambda_1 \neq \lambda_2).$$



The activity of the „2” isotope:

$$a_2(t) = \lambda_2 \cdot N_2(t) = a_1(0) \frac{\lambda_2}{\lambda_2 - \lambda_1} \left( e^{-\lambda_1 t} - e^{-\lambda_2 t} \right)$$

Using, that  $a_1(t) = a_1(0) \cdot e^{-\lambda_1 t}$  we get:

$$a_2(t) = a_1(t) \frac{\lambda_2}{\lambda_2 - \lambda_1} \left( 1 - e^{-(\lambda_2 - \lambda_1)t} \right)$$

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**Special cases:**

$$a_2(t) = a_1(t) \frac{\lambda_2}{\lambda_2 - \lambda_1} \left( 1 - e^{-(\lambda_2 - \lambda_1)t} \right)$$

1) If  $\lambda_2 > \lambda_1$ , then after long enough time the exponent gets negligibly small, and:

$$a_2(t) \approx a_1(t) \frac{\lambda_2}{\lambda_2 - \lambda_1}$$

From this we get:

Independent of time!

$$\frac{a_2(t)}{a_1(t)} = \frac{\lambda_2}{\lambda_2 - \lambda_1} = \text{konst.}$$

This is called **transient equilibrium**.

2) If  $\lambda_2 \gg \lambda_1$ , then the condition of transient equilibrium is met, but additionally,  $\lambda_1$  can be neglected in the denominator, and we get:

$$\frac{a_2(t)}{a_1(t)} = \frac{\lambda_2}{\lambda_2} = 1$$

With other words:  $a_1(t) = a_2(t)$

It can be shown similarly, that if  $\lambda_1$  is much smaller than all other decay constants, after long enough time we get

$$a_1(t) = a_2(t) = a_3(t) = \dots, \text{ in a many-member decay chain.}$$

This is called **secular equilibrium**.

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In secular equilibrium:  $a_1(t) = a_2(t) = a_3(t) = \dots$

Using  $a(t) = \lambda \cdot N(t) = \frac{N(t)}{T} \ln 2$  we get:

$$\frac{N_1(t)}{T_1} = \frac{N_2(t)}{T_2} = \frac{N_3(t)}{T_3} = \dots$$

This can be written in a different way:

$$N_1(t) : N_2(t) : N_3(t) \dots = T_1 : T_2 : T_3 : \dots$$

In secular equilibrium, the ratio of the quantities (number of atoms) of the members equals to the ratio of their half-lives.

Practical use: this makes possible the determination of very long half-lives (For example: the half life of  $^{238}\text{U}$  is 4,5 billion years.)

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## Radioactive dating

Using the decay properties of a radioactive substance one can conclude about the age of the sample

Problem: the initial conditions should be known!

Most commonly used isotopes for radioactive dating :

Isotope	Half life	Abundance
$^3\text{H}$ (tritium)	12,262 year	$1 \cdot 10^{-18}$
$^{14}\text{C}$ (radiocarbon)	5568 year	$2 \cdot 10^{-12}$
$^{40}\text{K}$	$1,3 \cdot 10^9$ year	$1,19 \cdot 10^{-4}$
$^{87}\text{Rb}$	$5 \cdot 10^{10}$ year	0,278
$^{238}\text{U}$	$4,51 \cdot 10^9$ year	0,992739
$^{235}\text{U}$	$7,04 \cdot 10^8$ year	0.007204
$^{232}\text{Th}$	$1,39 \cdot 10^{10}$ year	1.0

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For most precise dating the half-life of the isotope should be close to the age of the sample (at least similar orders of magnitude)

**Geological dating** (10 million years – few billion years)

- relative
- absolute

**Relative geological dating** (non-nuclear methods)

- paleontological (fossils in sedimentary rocks)
- based on the position in geological profile

**Absolute dating** (nuclear methods)

- Rubidium-strontium (Rb-Sr) method
- Lead-helium method (Th, or uranium decay chains)
- Potassium-argon method (K-Ar)

**Rubidium-strontium method:**  $^{87}\text{Rb} \xrightarrow{\beta, 50 \text{ billion y}} ^{87}\text{Sr}$

The age of the rock can be determined by the ratio:  $\frac{^{87}\text{Sr}}{^{87}\text{Rb}}$

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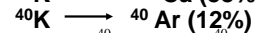
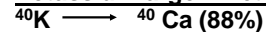
**Lead-helium method:** based on the radioactive decay chains.

- From  $^{238}\text{U}$  we get finally  $^{206}\text{Pb}$ . In the chain: 8  $\alpha$ -decays.
  - From  $^{232}\text{Th}$  we get  $^{208}\text{Pb}$ . In the chain: 6  $\alpha$ -decays.
  - From  $^{235}\text{U}$  we get  $^{207}\text{Pb}$ . In the chain: 7  $\alpha$ -decays
- Therefore helium and lead accumulate in the rock.

**Difficulties:**

- difficult to separate the lead-isotopes
- usually all the three decay chains are present
- the helium is a noble gas. Some quantity can escape
- all chains contain an isotope of Rn (radon) this is also a noble gas, can escape, the chain gets „broken”

**Potassium-argon method** ( $T = 1,3$  billion years)



The  $\frac{^{40}\text{Ca}}{^{40}\text{K}}$ , and  $\frac{^{40}\text{Ar}}{^{40}\text{K}}$  ratios should be measured

**Difficulties:**

- $^{40}\text{Ca}$  common, originates not only from the decay of  $^{40}\text{K}$
- $^{40}\text{Ar}$  is a noble gas, can escape.

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**Radiocarbon method** ( $T = 5568$  years)

The  $^{14}\text{C}$  is constantly produced in the atmosphere by the cosmic radiation.

Equilibrium abundance ( $\text{CO}_2$ ) in the air:  $^{14}\text{C}/^{12}\text{C} = 1,2 \cdot 10^{-12}$ .

Because of the continuous metabolism this concentration is present in the living creatures.

After the death, the metabolism stops, the feed of the  $^{14}\text{C}$  stops as well, it only decays.

The time elapsed since the death can be determined by measuring the radiocarbon concentration:

Here  $t$  is the time since the death,

$T$  is the half life of  $^{14}\text{C}$ .

$$\frac{N(^{14}\text{C})}{N(^{12}\text{C})} = 1,2 \cdot 10^{-12} \left( \frac{1}{2} \right)^{\frac{t}{T}}$$

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**Tritium method** ( $T = 12,26$  years)

The  $^3\text{H}$  is constantly produced in the atmosphere by the cosmic radiation.

Equilibrium abundance ( $\text{H}_2\text{O}$ ) in the air  $^3\text{H}/^1\text{H} = 1 \cdot 10^{-18}$ .

This concentration remains in the surface waters (rivers, lakes etc.) because of continuous exchange (rain).

The age of the underground water can be determined by measuring its tritium concentration.

Here  $t$  is the time since the water went underground,

$T$  is the half life of tritium.

$$\frac{N(^3\text{H})}{N(^1\text{H})} = 1 \cdot 10^{-18} \left(\frac{1}{2}\right)^{\frac{t}{T}}$$

**Note: age of dead living creatures cannot be determined, since the H-metabolism continues after the death!**

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**Self-test questions (contd.)**

9. Calculate the activity of 1 g  $^{226}\text{Ra}$ ! (Half-life: 1602 years)
10. An  $\alpha$ -counter detects 40% of the emitted  $\alpha$ -particles from a  $^{226}\text{Ra}$  sample. During one minute it counts 400.
  - a) What is the activity of the sample?
  - b) What is the precision of this value (in %)
  - c) What will be precision if we would measure 100 minutes?
11. Why do only 4 large decay-chains exist in nature?
12. Why are no  $\beta^+$ -decays and electron captures in the 4 large decay chains existing in nature?
13. In nuclear chain-reaction  $^{135}\text{I}$  and  $^{135}\text{Xe}$  isotopes are formed. When the chain reaction stops, they form a decay chain.  $^{135}\text{I} \rightarrow ^{135}\text{Xe} \rightarrow ^{135}\text{Cs}$ . The half-lives: 6,7 h ( $^{135}\text{I}$ ), 9,2 h ( $^{135}\text{Xe}$ )  
Determine the time-behaviour of the Xe isotope after the chain reaction stops. Will the number of Xe nuclei have a maximum? If yes, when?

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**Self-test questions (contd.)**

14. Pierre and Marie Curie extracted 1 g  $^{226}\text{Ra}$  from uranium ore. How much  $^{238}\text{U}$  was in the ore? (Suppose that their extraction efficiency was 100%!) Half-lives: 4,5 billion years ( $^{238}\text{U}$ ), 1602 years ( $^{226}\text{Ra}$ )
15. Now the abundance of  $^{235}\text{U}$  in natural uranium is 0,71%. Long time ago it was higher, reached also 5%. Then, the natural uranium could form a „natural nuclear reactor” with the ground water as moderator. (Remnants of such a natural nuclear reactors have been found in Oklo /South Africa/). How long ago could that happen? Half-lives: 4,5 billion years ( $^{238}\text{U}$ ), and 710 million years ( $^{235}\text{U}$ )
16. A wine-dealer argues that an old bottle of wine is 30 years old. Someone takes a small sample of the wine using a syringe through the cork, and finds that the  $^3\text{H}$  concentration is  $0,5 \cdot 10^{-18}$ . How old is the wine?

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